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MONTEZUMA SOLAR-CONSTANT
VALUES AND THEIR PERIODIC SOLAR
VARIATIONS

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MONTEZUMA SOLAR-CONSTANT VALUES AND THEIR PERIODIC SOLAR VARIATIONS

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We are convinced that solar-constant values from the Mount Montezuma, Chile, station are more accurate than those of any other Smithsonian station. This results from the meteorological superiority of the location. In three recent papers¹ (treating respectively of the 6.6456-day period in the solar radiation and in weather, of the trigger action of depressions of solar radiation to set off West Indian hurricanes, and of the effect of ionic bombardment of the earth to diminish solar radiation received here at times of great sunspot activity) I used the daily solar-constant values of Montezuma exclusively. The inclusion with them of less accurate data from our other stations would have been injurious in these studies of very small solar changes.

In volumes 5 and 6 of *Annals of the Smithsonian Astrophysical Observatory*, and in my paper "A Revised Analysis of Solar Constant Values"² the 10-day and monthly mean solar-constant values from several Smithsonian stations were combined in researches on long periods in solar variation. It seemed advisable to me to make a new search for long solar periodicities, using Montezuma data alone. I wished especially to test my former conclusion that all the periodic variations are integral submultiples of 273 months.

I have prepared a table of 10-day and monthly mean solar-constant values for Montezuma alone, from September 1923 to December 1947. They are given in table 1.

In table 1 the year and month are given in column 1. In column 2 appear the 10-day and monthly mean values of the solar constant, from Montezuma observations alone. Column 3 gives the number of days entering into these mean values. Readers should note that values in column 2 are to be understood as prefixed by the figures 1.9

¹ Smithsonian Misc. Coll., vol. 107, No. 4, 1947; vol. 110, Nos. 1 and 6, 1948.

² Smithsonian Misc. Coll., vol. 107, No. 10, 1947.

TABLE I.—Ten-day and monthly means, *Montezuma solar-constant values*
 Values given assumed to be prefixed by 1.9. Thus, 1.9536, etc.

1923		1924		1926		1927		1928	
9 I	536 9	12 I	506 5	3 I	439 7	6 I	479 8	9 I	— 0
II	531 8	II	470 1	II	515 4	II	427 6	II	442 4
III	540 8	III	518 5	III	393 8	III	433 9	III	410 4
M	536 25	M	508 11	M	435 19	M	447 23	M	426 8
10 I	446 9	1925		4 I	317 10	7 I	445 6	10 I	452 9
II	410 7	1 I	442 5	II	396 10	II	440 8	II	487 6
III	452 6	II	490 2	III	398 9	III	451 11	III	390 7
M	436 22	III	360 1	M	369 29	M	446 25	M	442 22
11 I	392 5	M	444 8	5 I	394 8	8 I	415 10	11 I	427 6
II	451 7	2 I	— 0	II	405 4	II	412 8	II	453 3
III	490 4	II	600 2	III	407 7	III	442 11	III	461 8
M	443 16	III	573 3	M	401 19	M	424 29	M	448 17
12 I	420 3	M	584 5	6 I	388 4	9 I	414 8	12 I	500 4
II	445 6	3 I	537 3	II	434 9	II	428 10	II	446 7
III	201 8	II	492 6	III	456 5	III	471 7	III	463 3
M	368 17	III	542 8	M	430 18	M	436 25	M	465 14
1924		M	524 17	7 I	439 10	10 I	481 7	1929	
1 I	416 7	4 I	536 8	II	424 9	II	438 5	1 I	450 1
II	443 8	II	530 8	III	433 11	III	417 9	II	590 3
III	459 9	III	477 3	M	432 30	M	443 21	III	485 2
M	441 24	M	524 19	8 I	472 8	11 I	450 8	M	532 6
2 I	369 8	5 I	463 8	II	433 3	II	446 5	2 I	432 4
II	460 1	II	484 9	III	475 6	III	436 5	II	390 3
III	422 6	III	479 8	M	466 17	M	445 18	III	335 4
M	396 15	M	476 25	9 I	441 9	12 I	479 7	M	385 11
3 I	523 9	6 I	420 4	II	396 5	II	388 4	3 I	443 3
II	380 7	II	496 5	III	457 8	III	381 7	II	360 9
III	434 7	III	480 5	M	437 22	M	421 18	III	422 5
M	453 23	M	469 14	10 I	373 9	1928		M	393 17
4 I	392 5	7 I	510 2	II	423 10	1 I	442 6	4 I	370 10
II	416 7	II	510 6	III	374 8	II	374 5	II	492 9
III	432 9	III	436 8	M	392 27	III	476 5	III	442 10
M	417 21	M	473 16	11 I	359 8	M	431 16	M	435 29
5 I	467 9	8 I	477 9	II	357 6	2 I	480 6	5 I	420 6
II	493 7	II	442 6	III	405 2	II	433 7	II	430 5
III	506 10	III	431 8	M	364 16	III	470 1	III	436 10
M	489 26	M	452 23	12 I	333 3	M	456 14	M	430 21
6 I	554 7	9 I	524 8	II	359 8	3 I	464 7	6 I	395 6
II	492 5	II	470 10	III	370 1	II	464 5	II	344 5
III	522 6	III	471 8	M	353 12	III	468 6	III	413 7
M	526 18	M	487 26	1927		M	466 18	M	388 18
7 I	511 7	10 I	452 8	1 I	396 9	4 I	430 8	7 I	397 8
II	544 10	II	500 6	II	357 3	II	411 8	II	407 9
III	470 9	III	458 9	III	375 2	III	408 6	III	420 9
M	510 26	M	467 23	M	385 14	M	417 22	M	409 26
8 I	542 5	11 I	420 6	2 I	348 4	5 I	436 9	8 I	396 8
II	411 8	II	487 10	II	467 7	II	511 8	II	398 6
III	390 5	III	470 6	III	— 0	III	468 9	III	402 6
M	442 18	M	464 22	M	424 11	M	470 26	M	399 20
9 I	462 5	12 I	482 10	3 I	— 0	6 I	472 9	9 I	397 6
II	483 7	II	460 3	II	500 6	II	475 4	II	381 9
III	431 8	III	497 7	III	466 9	III	460 3	III	416 9
M	457 20	M	484 20	M	479 15	M	471 16	M	397 24
10 I	536 5	1926		4 I	446 5	7 I	446 10	10 I	413 7
II	524 8	1 I	473 7	II	472 9	II	435 6	II	442 4
III	528 11	II	499 7	III	414 8	III	413 6	III	370 3
M	528 24	III	390 5	M	445 22	M	434 22	M	412 14
11 I	557 9	M	461 19	5 I	426 8	8 I	444 5	11 I	419 9
II	494 8	2 I	405 2	II	426 5	II	449 7	II	422 5
III	498 6	II	410 5	III	420 11	III	440 6	III	480 8
M	520 23	III	400 1	M	423 24	M	444 18	M	442 22
		M	408 8						

TABLE I.—Continued

1936		1937		1938		1939		1941	
3 I	304 9	6 I	472 8	9 I	462 6	12 I	375 6	3 I	575 4
II	340 8	II	451 7	II	448 8	II	473 7	II	545 4
III	412 9	III	443 3	III	479 10	III	445 9	III	540 9
M	352 26	M	460 18	M	464 24	M	435 22	M	549 17
4 I	417 4	7 I	— 0	10 I	470 8	1940		4 I	500 9
II	440 2	II	456 8	II	495 6	I	460 5	II	520 4
III	496 8	III	459 7	III	522 9	II	450 3	III	525 2
M	466 14	M	457 15	M	497 23	III	453 11	M	509 15
5 I	463 6	8 I	491 9	11 I	495 6	2 I	437 7	5 I	553 7
II	442 6	II	474 9	II	532 10	II	434 10	II	— 0
III	466 10	III	484 7	III	537 3	III	385 4	III	602 6
M	459 22	M	483 25	M	521 19	M	426 21	M	575 13
6 I	495 6	9 I	449 7	12 I	514 5	3 I	469 8	6 I	590 5
II	461 8	II	487 9	II	538 9	II	362 10	II	575 2
III	498 5	III	404 10	III	493 10	III	401 7	III	557 4
M	482 19	M	445 26	M	514 24	M	407 25	M	575 11
7 I	504 6	10 I	419 8	1939		4 I	435 2	7 I	637 3
II	441 7	II	434 10	I	— 0	II	458 9	II	560 9
III	455 11	III	540 5	II	— 0	III	544 7	III	554 5
M	463 24	M	452 23	III	400 2	M	489 18	M	572 17
8 I	420 4	11 I	503 9	2 I	424 7	5 I	519 9	8 I	566 7
II	470 6	II	457 6	II	— 6	II	486 8	II	542 4
III	455 8	III	505 6	III	463 6	III	510 9	III	480 6
M	442 18	M	490 21	M	442 13	M	506 26	M	529 17
9 I	385 4	12 I	518 5	3 I	420 3	6 I	486 10	9 I	540 8
II	430 10	II	570 3	II	442 5	II	491 7	II	509 7
III	486 9	III	541 8	III	452 8	III	516 7	III	585 2
M	444 23	M	540 16	M	443 16	M	496 24	M	532 17
10 I	426 5	1938		4 I	427 8	7 I	520 9	10 I	525 8
II	494 8	I	490 1	II	453 7	II	493 6	II	479 7
III	486 5	II	534 8	III	427 3	III	525 6	III	511 7
M	473 18	III	520 3	M	437 18	M	514 21	M	506 22
11 I	500 6	2 I	— 0	5 I	420 6	8 I	508 6	11 I	515 6
II	506 7	II	440 2	II	383 6	II	484 8	II	472 6
III	484 7	III	— 0	III	393 11	III	508 5	III	500 9
M	496 20	M	440 2	M	397 23	M	496 19	M	496 21
12 I	525 4	3 I	505 2	6 I	393 7	9 I	552 6	12 I	525 8
II	496 5	II	456 5	II	378 8	II	564 7	II	492 4
III	— 0	III	400 1	III	402 9	III	479 8	III	548 9
M	509 9	M	461 8	M	391 24	M	528 21	M	529 21
1937		4 I	445 8	7 I	420 8	10 I	450 2	1942	
I	— 0	II	432 6	II	388 9	II	472 4	I	565 8
II	450 1	III	462 6	III	398 6	III	486 7	II	536 5
III	477 4	M	446 20	M	402 23	M	476 13	III	494 5
M	472 5							M	537 18
2 I	513 9	5 I	458 9	8 I	373 8	11 I	408 6	2 I	496 7
II	470 4	II	430 5	II	354 5	II	447 9	II	499 7
III	516 5	III	396 7	III	417 8	III	414 7	III	460 6
M	504 18	M	430 21	M	385 21	M	426 22	M	486 20
3 I	389 8	6 I	420 6	9 I	476 7	12 I	429 7	3 I	413 3
II	403 6	II	454 5	II	447 6	II	492 5	II	409 10
III	407 8	III	460 4	III	452 8	III	507 4	III	444 5
M	400 22	M	442 15	M	459 21	M	468 16	M	419 18
4 I	372 9	7 I	441 8	10 I	447 6	1941		4 I	424 5
II	424 7	II	445 8	II	426 7	I	483 7	II	464 10
III	431 7	III	437 6	III	352 10	II	533 8	III	426 7
M	406 23	M	441 22	M	399 23	III	410 1	M	443 22
5 I	343 3	8 I	462 5	11 I	380 4	2 I	525 4	5 I	422 4
II	469 7	II	466 8	II	399 9	II	— 0	II	490 7
III	490 6	III	458 6	III	404 8	III	590 6	III	482 9
M	453 16	M	463 19	M	397 21	M	564 10	M	473 20

to give the complete solar constant in calories per square centimeter per minute.

Figure 1 shows graphically in curve A the march of the monthly mean values given in table 1. Curve B, on the same scale, gives departures from 1,945 calories remaining after 14 periodicities specified in table 2, below, have been removed from the original data given in column 2, table 1.

Table 2 also gives the yearly mean values, and numbers of days entering into them. It gives also smoothed-curve values derived from these yearly data, after plotting them as shown in figure 2. In the statistical search for periodic variations reported below, the smoothed-curve yearly mean values of table 2 were first to be removed by subtraction from the original monthly means. In order to do this the smoothed yearly means were first expanded graphically into a plot of smoothed monthly means. I do not take space to publish these smoothed monthly means, as their simple derivation will be easily understood, and as it makes no appreciable errors in the periodicities, to be given in table 2, whether these smoothed monthly means for eliminating yearly changes of the solar constant are the best that could be found or not; for these periodicities are found as means from statistical tables including many repetitions of the periods, and local errors are smoothed out.

In previous analysis of solar-constant values ³ numerous periodicities in solar variation were found to proceed simultaneously, all being approximately integral submultiples of 273 months in length. I did not wish to adopt this master period of 273 months in this present research without independently confirming it from Montezuma data alone. Figure 2, however, itself seems to indicate that a period of about this length would fit the yearly variations of the solar constant. There are researches of other authors which support the validity of a period approximating two 11-year sunspot cycles, as being in evidence in various solar and terrestrial phenomena. Thus G. E. Hale discovered that magnetism in sunspots reverses its polarity in a remarkable way with each successive sunspot cycle of 11 years, so that the sun's magnetic condition is restored only after two 11-year cycles pass, or about 22 $\frac{2}{3}$ years. A. E. Douglass has remarked a 23-year period in tree-ring widths. Various meteorologists have found it in terrestrial data. I myself pointed out that Wild's meteorological studies of the Russian Empire, when supplemented by later data, showed very clearly a 23-year cycle in weather at St. Petersburg.

³ Ann. Astrophys. Obs., vol. 6, p. 181, 1942; Smithsonian Misc. Coll., vol. 107, No. 10, 1947.

1960
58
56
54
52
50
48
46
44
42
40
38
1936
1948
46
44
42
40
38
1936



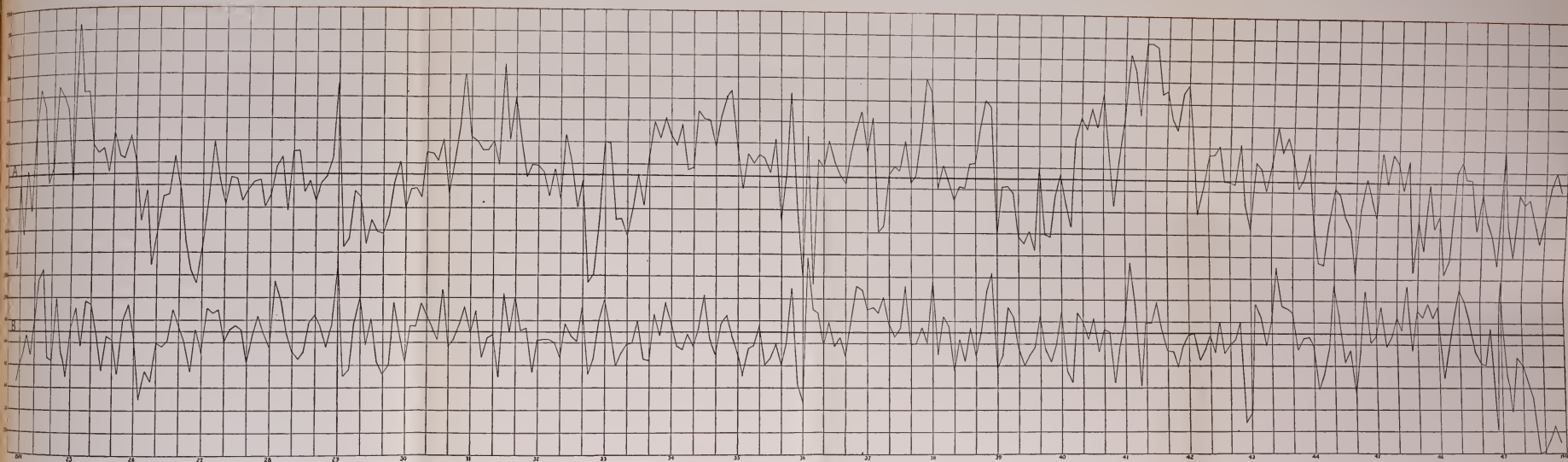


FIG. 1.—A, monthly mean march of solar constant, December 1923–December 1947, Montezuma station; B, residuals after periodic fluctuations removed.



Nevertheless, I began this present research without assuming a 273-month master period. First of all I removed the yearly variation from the values in column 2, table 1, as noted above. I then plotted the residual values and found that by far the most prominent periodic variation displayed in a large-scale plot of the residuals was of about 39 months. Seeking to fix its length as accurately as possible, by careful inspection of the large-scale plot, I finally decided on $39\frac{1}{2}$ months. I am not sure that the period may not be 39 months, which is exactly $1/7$ of 273 months; for the presence in the data of many other periodicities, and of accidental errors of observation, makes fixing of the exact length of a long period doubtful. Nevertheless, a table was prepared of seven columns, alternately of 39 and of 40 months in length. The mean of these columns is plotted in figure 3, *c*. As the reader will see, the march of this $39\frac{1}{2}$ -month periodicity is nearly a regular sine curve, and its amplitude is 0.0069 calorie, more than one-third of 1 percent of the solar constant.

The $39\frac{1}{2}$ -month periodicity was removed by subtraction to give a second list of monthly residuals. These also were plotted on a very large scale. There showed then a periodicity of considerable amplitude, approximately 91 months in length. A table 91 months long of three columns was made from the second residuals. With so few columns entering into the mean it seemed best to smooth the mean values by 5-month running means of them. The smoothed values being plotted, the 91-month periodicity appeared plainly, but superposed thereon there appeared a period of $\frac{1}{6}$ of 91 months. As it would be preferable to determine this curve of about 15 months by itself at a later stage, a smooth curve was drawn of 91-months period, cutting symmetrically through the 15-month superposed excrescences. The 91-month periodicity had the amplitude 0.0054 calorie. It is not of sine form, but rises rapidly to maximum, and falls slowly to minimum, like the well-known sunspot frequency curve of 11 years. This 91-month periodicity was removed from the data, leaving a third list of residuals, which were plotted on a large scale.

The third list, when plotted, showed clearly a strong periodic fluctuation of about 68 months. This was determined by forming a table of four columns, taking their mean, smoothing it by 5-month running means, and plotting the smoothed means in a curve given in figure 3, *b*. Very clearly there is a period of $1/7$ of 68 months superposed on the principal curve. Not wishing to evaluate a $9\frac{3}{4}$ -month periodicity until a later stage, I drew a smoothed curve as shown in figure 3, *b*. It is nearly of sine form, and has an amplitude of 0.0053 calorie, slightly under one-third of 1 percent of the solar constant.

It was now apparent from the behavior of the yearly variation of the solar constant, the excellence of the $39\frac{1}{2}$ -, 91-, and 68-month periodic curves, and the superposition of curves of $91/6$ and $68/7$ months, as noted above, that it is quite justified to regard 273 months as a master cycle in solar variation, and that many periodicities, nearly or exactly integral submultiples of 273 months, exist simultaneously therein. In all my subsequent search for periodicities in solar variation, as displayed in Montezuma solar-constant values, I accepted the 273-month master period, and sought for integral submultiples of it.

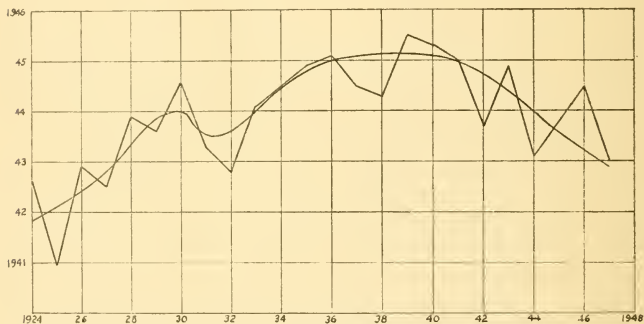


FIG. 2.—Yearly march of solar constant, 1924-1947.

Proceeding by the methods explained above, the periodicity of $54\frac{1}{2}$ months was next sought, found, and determined. Its amplitude is 0.0020 calorie, its form, like that of 91 months, comprises a rapid rise and slow fall. The curve, though smoothed by 5-month running means, has excrescences indicating the encroachment of a period approximating 8 months. Study of it was postponed, like those found with the 91- and 68-month periodicities, for later determination.

Attempts were then made to determine periodicities of $45\frac{1}{2}$, 34, and $30\frac{1}{3}$ months. But these proved so far dominated and obscured by variations of shorter periods that they were all passed over for the time. However the curve drawn when seeking a periodicity of $30\frac{1}{3}$ months clearly indicated a periodicity of half that length, of fairly large amplitude. So the next search made concerned $15\frac{1}{6}$ months. It will be noted that solar variations of 273, 91, 68, and $54\frac{1}{2}$ months period had now been extracted from the monthly data, and that the fourth list of residuals was now being used.

A period of $15\frac{1}{2}$ months is $1/18$ of 273 months. It was now practicable to divide the data into three groups, and tabulate them in 6-line tables of 15 columns.⁴ In this way it could be decided if the supposed $15\frac{1}{2}$ -month period continued in all three sections of the interval of 273 months. Figure 3, *a*, gives the mean curves for the three tabulations and the general mean. The three group means

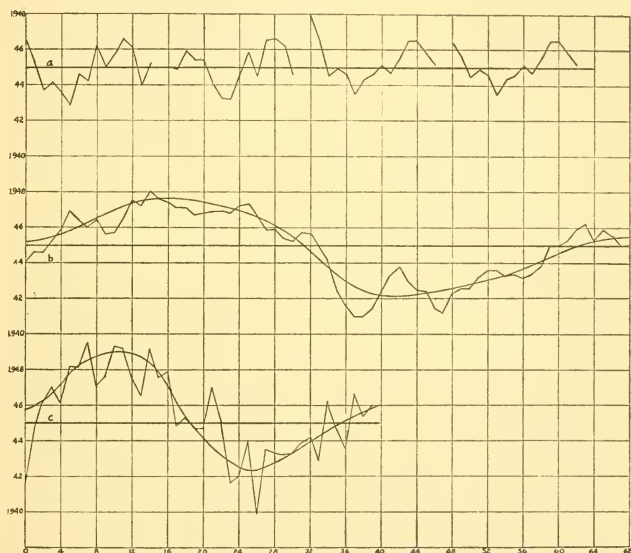


FIG. 3.—Examples of solar periodic fluctuations. *a*, $15\frac{1}{2}$ months. Observations 1924-31; 1932-39; 1940-47; and 1924-47; *b*, 68 months. 1924-47; *c*, $39\frac{1}{2}$ months. 1924-47.

show no certain secular displacement of maxima and minima, have nearly similar forms, and nearly equal amplitudes. Hence their mean was taken as shown in figure 3, *a*, and is regarded as a very well-determined periodicity of solar variation with an amplitude of 0.0030 calorie. This mean curve, being well supported in detail by the group means, is used unsmoothed, and the departures of it from

⁴ Whenever a periodicity not of exact months is determined, values or columns are omitted occasionally in tabulations, so that the mean values of columns fit the exact length of the periodicity.

1.945 calories were subtracted from the fourth list, giving a fifth list of residuals.

Though convinced of the validity of the assumption of a 273-month master cycle, I have passed over any discussion of the sunspot cycle of $11\frac{1}{2}$ years, approximating one-half of the master period. I now take up its consideration before noting the discovery of several other periodicities. Figure 2, which displays the variation of the yearly means of Montezuma solar-constant values, does, indeed, show depressions at the years 1925, 1931-32, 1937-38, and 1944. These may indicate a sunspot-cycle influence, but might better be attributed to the 68-month cycle which has already been discussed. Moreover, these depressions appearing in figure 2 are very small, with amplitudes only about $1/12$ of 1 percent of the solar constant, yet the 68-month curve, when specifically determined as given above, has an amplitude approaching $\frac{1}{3}$ of 1 percent.

Meteorologists recognize that the 11-year sunspot cycle is reflected in temperature, precipitation, and barometric pressure. Aldrich, also, has shown⁵ by the study of individual daily values of the sunspot numbers, and of solar-constant values, that there is a complex correlation between these phenomena. But my residual plots of monthly solar-constant values do not show any 136-month periodicity of appreciable amplitude. This is not really in contradiction to the findings of meteorologists. It is well known that the sunspot areas bombard the earth with electric ions. These, by acting as centers of condensation for water vapor and dust in the earth's atmosphere, may very well be competent to produce meteorological changes. Besides this, the ozone contents of the atmosphere may be affected by them in a way to influence meteorological phenomena. So we may recognize two kinds of solar influences on meteorology. One depends on variations of the solar radiation, the other on variations of ionic bombardment.

Having discovered and evaluated periodicities of 273, 91, 68, $54\frac{1}{2}$, $39\frac{1}{2}$, and $15\frac{1}{6}$ months in the variation of solar radiation, as evidenced by monthly mean solar-constant values of Montezuma, I next used the original 10-day mean values to seek for periodicities of less than 12 months. For such short periods the longer ones hitherto discussed produce no sensible interference. It would be tedious to recite all these trials. The method was always the same. By means of a long paper scale divided at regular intervals to represent a suspected period, I tested on the long plot of 10-day means whether such a

⁵ Smithsonian Misc. Coll., vol. 104, No. 12, 1945.

period seemed to be likely. If it seemed so, I arranged the 10-day mean values in groups of tables, each comprising about one-fourth of the total interval 1924-1947. They were never less than six lines long, and with as many columns as there were 10-day intervals in the proposed period. Where periods were not exact multiples of 10 days, values were omitted, or columns were omitted, occasionally, to bring the average lengths of the lines to that of the proposed period. The criterion of a true period was always that the several group tables agreed substantially in their means, as to phases and amplitudes of the suggested period, throughout the whole 273 months. Such good agreement is shown for the $15\frac{1}{6}$ -month period in figure 3, *a*. In several cases proposed periods failed to meet this test, and were rejected. Sometimes the phases shifted regularly from group to group through the 273-month interval. In such cases the period was shortened or lengthened to give unchanging phases.

As a result of this branch of the investigation, periodicities of $5\frac{2}{15}$, 8.035, $9\frac{3}{4}$, $11\frac{1}{3}$, $11\frac{15}{16}$ months were recognized as true, according to the above criterion. Being incommensurable in length, there was no need to subtract them one by one from the data. They could not materially influence each other. After determining them in the 10-day mean data, they were transformed into monthly means. Then their marches were tabulated throughout the 273 months, their amplitudes added algebraically at each month, and the algebraic total per month was subtracted from the fifth residual list, remaining after removing the longer periodicities named above. This left a sixth list of monthly residuals for further exploration.

To shorten a tedious story, the methods explained above, when applied to the sixth list of residuals, discovered additional periodicities of $14\frac{1}{3}$, $19\frac{1}{2}$, and $24\frac{1}{2}$ months. When all had been removed from the data, no other periodicities seemed worth investigation in the residual plot remaining. It is plotted as curve B of figure 1.⁶ The mean of the departures from 1.945 calories in curve B is 0.00189 calorie, or 0.097 percent of the solar constant. Many of the larger departures, which materially raise the mean as just given, occur in

⁶ One disturbing feature will be noted in figure 1, B. Though the year 1947 shows no remarkable eccentricity in curve A, it gives a great slump of $\frac{1}{2}$ percent in curve B. This is strange, for all the periodicities seem to fit the last year's data, including 1946, as well as the earlier years, as we see from figure 1, B. One notes, however, that curve A of figure 1 is almost entirely below 1.945 calories in 1947. It may be that the Montezuma values of 1947 are subject to a yet undiscovered error. Further observations of future years will decide.

the months December to February, when the atmospheric conditions at Montezuma are less favorable, and when many days are lost to observation. It cannot be claimed that the periodicities removed are perfectly correct in forms. Hence the final residuals are larger than they should be on this account also. We may conclude that of the variations of solar radiation indicated in figure 1,A, and which exceed 1 percent in range, accidental error of observation contributes less than 2/10 percent, and the periodic variations nearly 1 percent of the total range.

In table 2 I gave the details of the 14 periodicities in the variation of solar radiation, which have been discovered. There may be others of less than 5-months period, some of minor amplitude, and still others exceeding 273 months in period, which our observations have not yet continued long enough to discover. Indeed the large fluctuations of Great Lakes levels occurring at intervals of about 45 and 91 years seem to indicate that the double and quadruple of the master period of 273 months are of very great importance in meteorology. There is also the noted Bruckner period, of about three sunspot cycles, which may also be found eventually in solar-constant values if they continue to be observed for some years longer.