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MONTEZUMA SOLAR-CONSTANT VALUES AND THEIR PERIODIC SOLAR VARIATIONS

 $_{\rm BY}$

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MONTEZUMA SOLAR-CONSTANT VALUES AND THEIR PERIODIC SOLAR VARIATIONS

By C. G. ABBOT

Research Associate, Smithsonian Institution

We are convinced that solar-constant values from the Mount Montezuma, Chile, station are more accurate than those of any other Smithsonian station. This results from the meteorological superiority of the location. In three recent papers ¹ (treating respectively of the 6.6456-day period in the solar radiation and in weather, of the trigger action of depressions of solar radiation to set off West Indian hurricanes, and of the effect of ionic bombardment of the earth to diminish solar radiation received here at times of great sunspot activity) I used the daily solar-constant values of Montezuma exclusively. The inclusion with them of less accurate data from our other stations would have been injurious in these studies of very small solar changes.

In volumes 5 and 6 of Annals of the Smithsonian Astrophysical Observatory, and in my paper "A Revised Analysis of Solar Constant Values" ² the Io-day and monthly mean solar-constant values from several Smithsonian stations were combined in researches on long periods in solar variation. It seemed advisable to me to make a new search for long solar periodicities, using Montezuma data alone. I wished especially to test my former conclusion that all the periodic variations are integral submultiples of 273 months.

I have prepared a table of 10-day and monthly mean solar-constant values for Montezuma alone, from September 1923 to December 1947. They are given in table 1.

In table 1 the year and month are given in column 1. In column 2 appear the 10-day and monthly mean values of the solar constant, from Montezuma observations alone. Column 3 gives the number of days entering into these mean values. Readers should note that values in column 2 are to be understood as prefixed by the figures 1.9

¹ Smithsonian Misc. Coll., vol. 107, No. 4, 1947; vol. 110, Nos. 1 and 6, 1948.

² Smithsonian Misc. Coll., vol. 107, No. 10, 1947.

TABLE I.--Ten-day and monthly means, Montezuma solar-constant values Values given assumed to be prefixed by 1.9. Thus, 1.9536, etc.

	9	-												
1923 9 I 11 111 M	536 531 540 536	9 8 25	1924 12 I II III M	506 470 518 508	5 1 5 1 1	1926 3 I II III M	439 515 393 435	7 4 8 19	¹⁹²⁷ 6 I II III M	479 427 433 447	8 6 9 23	1928 9 I 11 111 M	442 410 426	0 4 4 8
io I III III M	446 410 452 436	9 7 6 22	1925 I I II III M	442 490 360 444	5 2 1 8	4 I III M	317 396 398 369	10 10 9 29	7 I II III M	445 440 451 446	6 8 11 25	10 I II III M	452 487 390 442	9 6 7 22
II I III M	392 451 490 443	5 7 4 16	² I II III M	600 573 584	0 2 3 5	5 I III M	394 405 407 401	8 4 7 19	8 I II III M	415 412 442 424	10 8 11 29	II II III M	427 453 461 448	6 3 8 17
12 I II III M 1924_	420 445 291 368	3 6 8 17	³ II III M	537 492 542 524	3 6 8 17	6 I III M	388 434 456 430	4 9 5 18	9 I II III M	414 428 471 436	8 10 7 25	12 I II III M 1929	500 446 463 465	4 7 3 14
IJ24 II III M	416 443 459 441	7 8 9 24	4 I II M	536 530 477 524	8 8 3 19	7 I II III M	439 424 433 432	10 9 11 30	10 I II III M	481 438 417 443	7 5 9 21	1929 I I II III M	450 590 485 532	1 3 2 6
² I JII M	369 460 422 396	8 1 6 15	5 I II III M	463 484 479 476	8 9 8 25	8 I II III M	472 433 475 466	8 3 6 17	II II III M	450 446 436 445	8 5 18	² I II III M	432 390 335 385	4 3 4 11
³ II III M	523 380 434 4 53	9 7 23	6 I II III M	420 496 480 469	4 5 5 14	9 I II III M	441 396 457 437	9 5 22	12 I II III M 1928	479 388 381 421	7 4 7 18	3 I II III M	443 360 422 393	3 9 5 17
4 I III M	392 416 432 417	5 7 9 21	7 I II III M	510 510 436 473	2 6 8 16	IO I II III M	373 423 374 392	9 10 8 27	I I II III M	442 374 476 431	6 5 5 16	${}^{4}_{\mathrm{II}}{{}^{\mathrm{II}}_{\mathrm{III}}}_{\mathrm{M}}$	370 492 442 435	10 9 10 29
5 I III III M	467 493 506 489	9 7 10 26	8 I II III M	477 442 431 452	9 6 8 23	II I II III M	359 357 405 364	8 6 2 16	² I II III M	480 433 470 456	6 7 1 14	5 I II III M	420 430 436 430	6 5 10 21
6 I II III M	554 492 522 526	7 5 18	9 I II III M	524 470 471 487	8 10 8 26	12 I II III M	333 359 370 353	3 8 1 12	3 I II III M	464 464 468 466	7 5 18	6 I II III M	395 344 413 388	6 5 7 18
7 I II III M	511 544 470 510	7 10 9 26	10 I II III M	452 500 458 467	8 6 9 23	1927 1 I 11 111 M	396 357 375 385	9 3 2 14	4 I II III M	430 411 408 417	8 8 6 22	7 I II III M	397 407 420 409	8 9 9 26
8 I III M	542 411 390 442	58 58 18	II I II III M	420 487 470 464	6 10 6 22	² I II III M	348 467 424	4 7 0 11	5 I II III M	436 511 468 470	9 8 9 26	8 I II III M	396 398 402 399	8 6 20
9 I III M	462 483 431 457	5 7 20	12 I II III M 1926	482 460 497 484	10 3 7 20	³ II III M	500 466 479	0 6 9 15	6 I II III M	472 475 460 471	9 4 3 16	9 I II III M	397 381 416 397	6 9 9 24
IO I II III M	536 524 528 528	5 8 11 24	I I III M	473 499 390 461	7 7 5 19	4 I III M	446 472 414 445	5 9 8 22	7 I II III M	446 435 413 434	10 6 6 22	IO I II III M	413 442 370 412	7 4 3 14
II I III M	557 494 498 520	9 8 6 23	² I II III M	405 410 400 408	2 5 1 8	5 I III M	426 426 420 423	8 5 11 24	8 I II III M	444 449 440 444	5 7 6 18	II I II III M	419 422 480 442	9 5 8 22

2

TABLE I.-Continued

1929 12 I II III M	468 477 426 461	58 58	1931 3 I 11 111 M	390 468 514 472	4 9 9 22	¹⁹³² 6 I II III M	446 500 505 486	5 6 17	1933 9 I 11 111 M	473 450 504 470	7 9 5 21	1934 12 I II III M	535 522 520 527	6 8 2 16
1930 I I III III M	460 412 395 420	2 4 2 8	${}^{4}_{\mathrm{II}}{{}^{\mathrm{II}}_{\mathrm{III}}}_{\mathrm{M}}$	520 462 432 472	5 4 5 14	7 I II III M	532 390 435 462	5 1 10 16	10 I II III M	493 500 499 498	3 7 7 17	¹⁹³⁵ 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1	482 460 467 475	9 2 4 15
² I II III M	398 460 480 437	6 1 5 12	5 I III M	520 460 480	0 1 2 3	8 I II III M	456 412 382 421	7 5 5 17	II I III M	438 509 510 484	5 8 1 14	² I II III M	455 430 410 439	4 1 2 7
³ II III M	450 422 443 43 ⁸	7 8 10 25	6 I II III M	473 480 432 459	7 2 5 14	9 I II III M	432 447 456 445	5 6 5 16	12 I II III M 1934 I I	512 480 505 502	5 4 10 19	3 I III III M	437 492 470	4 0 6 10
4 I II III M	444 401 443 4 30	8 7 6 21	7 I II III M	570 593 505 549	2 3 4 9	10 I III III M	414 310 346 354	56 56	II II III M	504 450 515 486	5 4 2 1 1	4 I III M	437 488 475 461	9 5 6 20
5 I II III M	450 490 470 470	4 4 12	8 I II III M	485 512 456 482	4 6 7 17	II II III M	370 342 377 361	7 8 6 21	² I II III M	555 455 462 477	2 2 6 10	5 I II III M	478 451 478 469	10 9 9 28
6 I II III M	463 462 483 469	3 4 3 10	9 I III III M	502 518 537 520	6 5 7 18	12 I II III M 1933_	373 457 3 ⁸ 7 411	3 4 3 10	³ I II III M	430 478 543 496	3 5 6 14	6 I II III M	463 484 455 466	9 5 20
7 I II III M	437 490 517 462	8 1 3 12	10 I II III M	544 450 452 482	7 6 8 21	II II III M	510 472 485 480	1 5 2 8	4 I II III M	503 456 435 455	3 8 8 19	7 I II III M	453 460 430 453	9 3 1 13
8 I II III M	473 490 479 481	3 4 11 18	II II III M	431 452 460 448	7 8 7 22	² I II III M	477 480 479	3 4 0 7	5 I II III M	480 466 440 457	6 5 11 22	8 I III M	484 514 460 483	7 10 9 26
9 I III III M	456 340 422 433	10 2 4 16	12 I II III M 1932_	444 477 459	8 7 0 15	3 I II III M	450 373 407 410	3 3 4 10	6 I II III M	518 516 460 508	6 5 2 13	9 I II III M	420 419 405 417	3 8 2 13
10 I II III M	453 470 462 461	9 6 8 23	II II III M	462 447 465 458	4 3 2 9	4 I III M	410 457 370 411	8 9 10 27	7 I II III M	524 502 477 501	7 6 7 20	IO I II III M	407 469 449 451	3 8 7 18
II I III M	482 475 528 492	5.00 5.00 I	² I II III M	435 492 432 453	4 4 12	5 I II III M	384 401 414 397	10 9 5 24	8 I II III M	507 497 499 500	4 6 7 17	II I II M	515 539 516 526	2 8 8 18
12 I II III M 1931	540 535 550 540	9 10 4 23	3 I II III M	363 447 444 431	3 7 7 17	6 I II III M	421 426 412 420	8 9 25	9 I II III M	474 490 466 477	10 9 7 26	12 I II III M 1936	415 437 400 421	4 3 2 8
I I II III M	430 497 484	і б 7	4 I III III M	497 441 440 455	4 9 3 16	7 I II III M	415 471 449 451	4 8 7 19	IO I II III M	495 497 515 503	8 7 23	II III M	378 340 300 361	6 2 1 9
² I II III M	4 ⁸ 5 497 458 4 80	2 6 5 13	5 I III III M	452 410 429	4 5 0 9	8 I III M	428 420 423	0 6 9 15	II I III M	530 528 504 520	9 5 22	² I II III M	500 492 440 4 ⁸ 7	5 5 2 1 2

TABLE I.—Continued

1936 3 I II III M	304 340 412 352	9 8 9 26	¹⁹³⁷ 6 I II III M	472 451 443 460	8 7 3	1938 9 I 11 111 M	462 448 479 464	6 8 10 24	1939 12 I II III M 1940_	375 473 445 435	6 7 9 22	¹⁹⁴¹ 3 I II III M	575 545 540 549	4 4 9 17
4 I III M	417 440 496 466	4 2 8 14	7 I III M	456 459 457	0 8 7 15	10 I III M	470 495 522 497	8 6 9 23	II II III M	460 450 453 454	5 3 11 19	4 I II III M	500 520 525 509	9 4 2 15
5 I III III M	463 442 466 459	6 6 10 22	8 I III M	49 I 474 484 483	9 9 7 25	II I III M	495 532 537 521	6 10 3 19	² I II III M	437 434 385 426	7 10 4 21	5 I II III M	553 602 575	7 0 6 13
6 I II III M	495 461 498 482	6 8 5 19	9 I II III M	449 487 404 445	7 9 10 26	12 I II III M 1939 I I	514 538 493 514	5 9 10 24	3 I II III M	469 362 401 407	8 10 7 25	6 I II III M	590 575 557 575	5 2 4 11
7 I III III M	504 441 455 463	6 7 11 24	10 I II III M	419 434 540 452	8 10 5 23	II II III M	400 400	0 0 2 2	4 I II III M	435 458 544 489	2 9 7 18	7 I II III M	637 560 554 572	3 9 5 17
8 I III M	420 470 455 452	4 8 18	II I III III M	503 457 505 490	9 6 6 21	² I II III M	424 463 442	7 0 6 13	5 I II III M	519 486 510 506	9 8 9 26	8 I II III M	566 542 480 529	7 4 6 17
9 I III III M	385 430 486 444	4 10 9 23	12 I II III M 1938_	518 570 541 540	5 3 16	3 I II III M	420 442 452 443	3 5 8 16	6 I II III M	486 491 516 496	10 7 7 24	9 I II III M	540 509 585 532	8 7 2 17
IO I II III M	426 494 486 473	58 58 18	II II III M	490 534 520 527	1 8 3 12	4 I II III M	427 453 427 437	8 7 18	7 I II III M	520 493 525 514	9 6 21	10 I II III M	525 479 511 506	8 7 22
II I III M	500 506 484 496	6 7 7 20	² I II III M	440 440	0 2 0 2	5 I II III M	420 383 393 397	6 6 11 23	8 I II III M	508 484 508 496	6 8 5 19	II II III M	515 472 500 496	6 6 9 21
12 I II III M 1937_	525 496 509	4 5 9	3 I II III M	505 456 400 461	2518	6 I II III M	393 378 402 391	7 8 9 24	9 I II III M	552 564 479 528	6 7 8 21	12 I II III M 1942	525 492 548 529	8 4 9 21
II III M	450 477 472	0 1 4 5	${}^{4}_{{}^{1}II}_{{}^{1}II}_{{}^{1}M}$	445 432 462 446	8 6 20	7 I II III M	420 388 398 402	8 9 6 23	IO I II III M	450 472 486 476	2 4 7 13	I I II III M	565 536 494 537	8 5 5 18
² I II III M	513 470 516 504	9 4 5 18	5 I III III M	458 430 396 430	9 5 7 21	8 I III III M	373 354 417 385	8 5 21	II II III M	408 447 414 426	6 9 7 22	² I II III M	496 499 460 486	7 7 6 20
3 I III III M	389 403 407 400	8 6 8 22	6 I II III M	420 454 460 442	6 5 4 15	9 I II III M	476 447 452 459	7 6 8 21	12 I II III M 1941	429 492 507 468	7 5 4 16	3 I II III M	413 409 444 419	3 10 5 18
4 I III M	372 424 431 406	9 7 7 23	7 I III M	441 445 437 441	8 6 22	IO I II III M	447 426 352 399	6 7 10 23	II II III M	483 533 410 504	7 8 1 16	4 I II III M	424 464 426 443	5 10 7 22
5 I III III M	343 469 490 453	3 7 6 16	8 I II III M	462 466 458 463	5 8 6 19	II II III M	380 399 404 397	4 9 8 21	² I II III M	525 590 564	4 0 6 10	5 I III M,	422 490 482 473	4 7 9 20

TABLE 1.—Continued

¹⁹⁴² 6 I II III M	4 ⁸ 3 475 464 474	7 6 7 20	¹⁹⁴³ 8 I II III M	469 485 512 490	10 4 11 25	1944 10 I II III M	468 432 364 420	10 6 8 24	¹⁹⁴⁵ 12 I II III M	421 409 391 407	8 7 8 23	¹⁹⁴⁷ ² I II III M	360 410 4 20 409	1 3 4 8
7 I II III M	494 490 464 482	7 7 8 22	9 I II III M	470 476 469 472	9 9 7 25	II II III M	455 436 470 453	8 9 8 25	1946 II II M	375 410 489 418	10 7 7 24	³ II III M	342 382 413 381	9 8 11 28
8 I II III M	450 461 436 449	8 9 10 27	IO I II III M	447 462 430 443	8 5 11 24	12 I II III M 1945	422 410 446 435	5 1 8 14	² I II III M	378 360 357 365	5 3 7 15	4 I II III M	380 450 458 439	3 6 15
9 I II III M	451 443 450 449	8 3 8 19	II II III M	478 390 410 453	6 1 2 9	II II III M	420 413 440 417	2 7 1 10	3 I II III M	381 363 393 378	7 8 7 22	5 I II III M	450 443 411 430	2 7 7 16
10 I II III M	443 440 456 447	8 5 9 22	I2 I II III M 1944	450 415 502 475	1 4 10 15	² II III M	503 43 5 480 477	9 6 2 17	4 I II III M	492 488 370 459	9 5 5 19	6 I II III M	384 470 454 435	9 8 9 26
II II III M	439 485 526 483	8 10 8 26	1944 II III M	403 440 412	3 0 1 4	3 I II III M	451 458 415 448	8 10 4 22	5 I II III M	486 486 433 469	5 8 6 19	7 I II III M	428 378 424 414	9 5 5 19
12 I II III M 1943_	429 404 447 428	10 8 10 28	² I II III M	460 347 375	1 3 0 4	4 I II III M	451 488 487 475	10 10 10 30	6 I II III M	432 470 458 455	4 5 5 14	8 I II III M	373 387 416 393	4 7 5 16
II II III M	404 419 390 406	9 9 6 24	³ II III M	377 377 368 373	6 4 18	5 I II III M	469 453 481 467	10 10 8 28	7 I II III M	484 495 406 455	5 6 8 19	9 I II III M	422 432 395 418	6 5 4 15
² I II III M	465 486 443 467	2 8 6 16	4 I III M	373 434 427 413	6 8 6 20	6 I II III M	466 436 418 442	9 10 6 25	8 I II III M	409 396 420 405	8 7 2 17	IO I II III M	429 433 479 445	9 7 7 23
³ II III M	425 501 439 461	4 8 20	5 I II III M	457 442 431 443	8 5 21	7 I II III M	464 472 475 470	8 9 6 23	9 I II III M	444 450 415 439	7 6 4 17	II II III M	479 454 438 460	8 8 4 20
4 I II III M	427 443 459 441	10 9 7 26	6 I II III M	441 430 443 438	9 10 10 29	8 I III M	421 417 243 366	7 7 6 20	IO I II III M	388 421 428 415	5 7 7 19	I2 I II III M	433 454 435 442	4 5 4 13
5 I II III M	463 467 465 465	7 10 4 21	7 I II III M	400 444 393 417	5 9 7 21	9 II III M	398 427 417 414	6 6 7 19	II II III M	437 395 270 400	4 4 1 9			
6 I II III M	503 497 498 500	6 4 5 15	8 I II III M	452 391 363 407	8 7 6 21	IO I II III M	441 391 324 386	7 9 7 23	12 I II III M 1947	300 398 373	1 0 5 6			
7 I II III M	484 443 496 476	7 7 9 23	9 I III III M	395 347 316 365	10 3 5 18	II II III M	486 409 465 448	5 9 10 24	1947 II III M	492 450 473 480	4 1 3			

to give the complete solar constant in calories per square centimeter per minute.

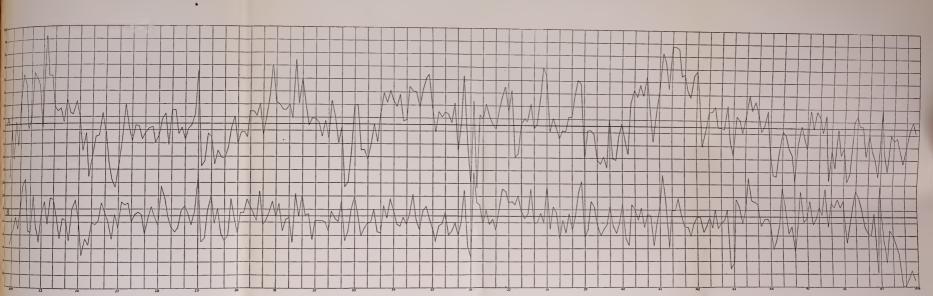
Figure I shows graphically in curve A the march of the monthly mean values given in table I. Curve B, on the same scale, gives departures from 1.945 calories remaining after 14 periodicities specified in table 2, below, have been removed from the original data given in column 2, table I.

Table 2 also gives the yearly mean values, and numbers of days entering into them. It gives also smoothed-curve values derived from these yearly data, after plotting them as shown in figure 2. In the statistical search for periodic variations reported below, the smoothedcurve yearly mean values of table 2 were first to be removed by subtraction from the original monthly means. In order to do this the smoothed yearly means were first expanded graphically into a plot of smoothed monthly means. I do not take space to publish these smoothed monthly means, as their simple derivation will be easily understood, and as it makes no appreciable errors in the periodicities, to be given in table 2, whether these smoothed monthly means for eliminating yearly changes of the solar constant are the best that could be found or not; for these periodicities are found as means from statistical tables including many repetitions of the periods, and local errors are smoothed out.

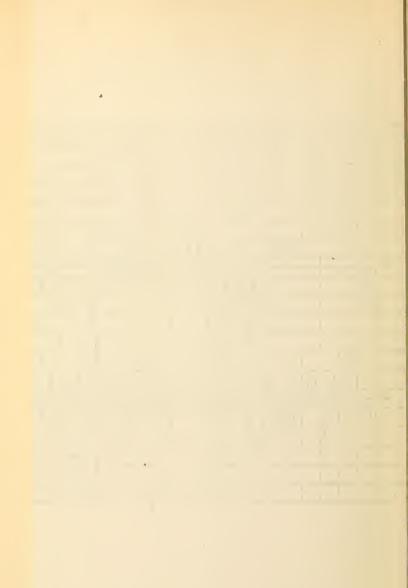
In previous analysis of solar-constant values ⁸ numerous periodicities in solar variation were found to proceed simultaneously, all being approximately integral submultiples of 273 months in length. I did not wish to adopt this master period of 273 months in this present research without independently confirming it from Montezuma data alone. Figure 2, however, itself seems to indicate that a period of about this length would fit the yearly variations of the solar constant. There are researches of other authors which support the validity of a period approximating two II-year sunspot cycles, as being in evidence in various solar and terrestrial phenomena. Thus G. E. Hale discovered that magnetism in sunspots reverses its polarity in a remarkable way with each successive sunspot cycle of II years, so that the sun's magnetic condition is restored only after two II-year cycles pass, or about 22²/₃ years. A. E. Douglass has remarked a 23-year period in tree-ring widths. Various meteorologists have found it in terrestrial data. I myself pointed out that Wild's meteorological studies of the Russian Empire, when supplemented by later data, showed very clearly a 23-year cycle in weather at St. Petersburg.

⁸ Ann. Astrophys. Obs., vol. 6, p. 181, 1942; Smithsonian Misc. Coll., vol. 107, No. 10, 1947.





F16. 1.--A, monthly mean march of solar constant, December 1923-December 1947, Montezuma station; B, residuals after periodic fluctuations removed.

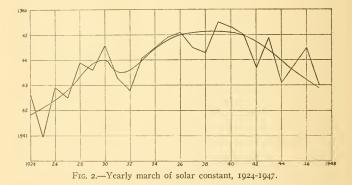


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Nevertheless, I began this present research without assuming a 273-month master period. First of all I removed the yearly variation from the values in column 2, table I, as noted above. I then plotted the residual values and found that by far the most prominent periodic variation displayed in a large-scale plot of the residuals was of about 30 months. Seeking to fix its length as accurately as possible, by careful inspection of the large-scale plot, I finally decided on 30¹ months. I am not sure that the period may not be 30 months, which is exactly 1/7 of 273 months; for the presence in the data of many other periodicities, and of accidental errors of observation, makes fixing of the exact length of a long period doubtful. Nevertheless, a table was prepared of seven columns, alternately of 30 and of 40 months in length. The mean of these columns is plotted in figure 3. c. As the reader will see, the march of this 39¹/₂-month periodicity is nearly a regular sine curve, and its amplitude is 0.0060 calorie, more than one-third of I percent of the solar constant.

The 39¹-month periodicity was removed by subtraction to give a second list of monthly residuals. These also were plotted on a very large scale. There showed then a periodicity of considerable amplitude, approximately of months in length. A table of months long of three columns was made from the second residuals. With so few columns entering into the mean it seemed best to smooth the mean values by 5-month running means of them. The smoothed values being plotted, the 91-month periodicity appeared plainly, but superposed thereon there appeared a period of 1 of 91 months. As it would be preferable to determine this curve of about 15 months by itself at a later stage, a smooth curve was drawn of 91-months period, cutting symmetrically through the 15-month superposed excrescences. The 91-month periodicity had the amplitude 0.0054 calorie. It is not of sine form, but rises rapidly to maximum, and falls slowly to minimum, like the well-known sunspot frequency curve of II years. This or-month periodicity was removed from the data, leaving a third list of residuals, which were plotted on a large scale.

The third list, when plotted, showed clearly a strong periodic fluctuation of about 68 months. This was determined by forming a table of four columns, taking their mean, smoothing it by 5-month running means, and plotting the smoothed means in a curve given in figure 3, b. Very clearly there is a period of I/7 of 68 months superposed on the principal curve. Not wishing to evaluate a $9\frac{3}{4}$ -month periodicity until a later stage, I drew a smoothed curve as shown in figure 3, b. It is nearly of sine form, and has an amplitude of 0.0053 calorie, slightly under one-third of I percent of the solar constant. It was now apparent from the behavior of the yearly variation of the solar constant, the excellence of the 39½-, 91-, and 68-month periodic curves, and the superposition of curves of 91/6 and 68/7 months, as noted above, that it is quite justified to regard 273 months as a master cycle in solar variation, and that many periodicities, nearly or exactly integral submultiples of 273 months, exist simultaneously therein. In all my subsequent search for periodicities in solar variation, as displayed in Montezuma solar-constant values, I accepted the 273-month master period, and sought for integral submultiples of it.



Proceeding by the methods explained above, the periodicity of $54\frac{1}{2}$ months was next sought, found, and determined. Its amplitude is 0.0020 calorie, its form, like that of 91 months, comprises a rapid rise and slow fall. The curve, though smoothed by 5-month running means, has excrescences indicating the encroachment of a period approximating 8 months. Study of it was postponed, like those found with the 91- and 68-month periodicities, for later determination.

Attempts were then made to determine periodicities of $45\frac{1}{2}$, 34, and $30\frac{1}{3}$ months. But these proved so far dominated and obscured by variations of shorter periods that they were all passed over for the time. However the curve drawn when seeking a periodicity of $30\frac{1}{3}$ months clearly indicated a periodicity of half that length, of fairly large amplitude. So the next search made concerned $15\frac{1}{6}$ months. It will be noted that solar variations of 273, 91, 68, and $54\frac{1}{2}$ months period had now been extracted from the monthly data, and that the fourth list of residuals was now being used.

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A period of $15\frac{1}{5}$ months is 1/18 of 273 months. It was now practicable to divide the data into three groups, and tabulate them in 6-line tables of 15 columns.⁴ In this way it could be decided if the supposed $15\frac{1}{5}$ -month period continued in all three sections of the interval of 273 months. Figure 3, *a*, gives the mean curves for the three tabulations and the general mean. The three group means

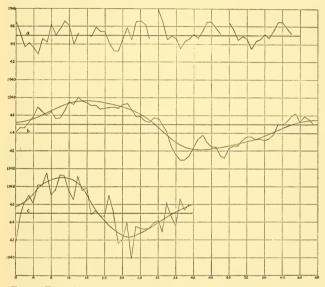


FIG. 3.—Examples of solar periodic fluctuations. a, $15\frac{1}{6}$ months. Observations 1924-31; 1932-39; 1940-47; and 1924-47; b, 68 months. 1924-47; c, $39\frac{1}{2}$ months. 1924-47.

show no certain secular displacement of maxima and minima, have nearly similar forms, and nearly equal amplitudes. Hence their mean was taken as shown in figure 3, *a*, and is regarded as a very welldetermined periodicity of solar variation with an amplitude of 0.0030 calorie. This mean curve, being well supported in detail by the group means, is used unsmoothed, and the departures of it from

9

⁴Whenever a periodicity not of exact months is determined, values or columns are omitted occasionally in tabulations, so that the mean values of columns fit the exact length of the periodicity.

1.945 calories were subtracted from the fourth list, giving a fifth list of residuals.

Though convinced of the validity of the assumption of a 273month master cycle, I have passed over any discussion of the sunspot cycle of $11\frac{1}{3}$ years, approximating one-half of the master period. I now take up its consideration before noting the discovery of several other periodicities. Figure 2, which displays the variation of the yearly means of Montezuma solar-constant values, does, indeed, show depressions at the years 1925, 1931-32, 1937-38, and 1944. These may indicate a sunspot-cycle influence, but might better be attributed to the 68-month cycle which has already been discussed. Moreover, these depressions appearing in figure 2 are very small, with amplitudes only about 1/12 of 1 percent of the solar constant, yet the 68month curve, when specifically determined as given above, has an amplitude approaching $\frac{1}{3}$ of 1 percent.

Meteorologists recognize that the II-year sunspot cycle is reflected in temperature, precipitation, and barometric pressure. Aldrich, also, has shown⁵ by the study of individual daily values of the sunspot numbers, and of solar-constant values, that there is a complex correlation between these phenomena. But my residual plots of monthly solar-constant values do not show any 136-month periodicity of appreciable amplitude. This is not really in contradiction to the findings of meteorologists. It is well known that the sunspot areas bombard the earth with electric ions. These, by acting as centers of condensation for water vapor and dust in the earth's atmosphere, may very well be competent to produce meteorological changes. Besides this, the ozone contents of the atmosphere may be affected by them in a way to influence meteorological phenomena. So we may recognize two kinds of solar influences on meteorology. One depends on variations of the solar radiation, the other on variations of ionic bombardment.

Having discovered and evaluated periodicities of 273, 91, 68, 54 $\frac{1}{2}$, $39\frac{1}{2}$, and $15\frac{1}{6}$ months in the variation of solar radiation, as evidenced by monthly mean solar-constant values of Montezuma, I next used the original 10-day mean values to seek for periodicities of less than 12 months. For such short periods the longer ones hitherto discussed produce no sensible interference. It would be tedious to recite all these trials. The method was always the same. By means of a long paper scale divided at regular intervals to represent a suspected period, I tested on the long plot of 10-day means whether such a

⁵ Smithsonian Misc. Coll., vol. 104, No. 12, 1945.

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period seemed to be likely. If it seemed so, I arranged the Io-day mean values in groups of tables, each comprising about one-fourth of the total interval 1924-1947. They were never less than six lines long, and with as many columns as there were Io-day intervals in the proposed period. Where periods were not exact multiples of IO days, values were omitted, or columns were omitted, occasionally, to bring the average lengths of the lines to that of the proposed period. The criterion of a true period was always that the several group tables agreed substantially in their means, as to phases and amplitudes of the suggested period, throughout the whole 273 months. Such good agreement is shown for the 15th-month period in figure 3, *a*. In several cases proposed periods failed to meet this test, and were rejected. Sometimes the phases shifted regularly from group to group through the 273-month interval. In such cases the period was shortened or lengthened to give unchanging phases.

As a result of this branch of the investigation, periodicities of 5 2/15, 8.035, $9\frac{3}{4}$, $11\frac{1}{3}$, 11 15/16 months were recognized as true, according to the above criterion. Being incommensurable in length, there was no need to subtract them one by one from the data. They could not materially influence each other. After determining them in the IO-day mean data, they were transformed into monthly means. Then their marches were tabulated throughout the 273 months, their amplitudes added algebraically at each month, and the algebraic total per month was subtracted from the fifth residual list, remaining after removing the longer periodicities named above. This left a sixth list of monthly residuals for further exploration.

To shorten a tedious story, the methods explained above, when applied to the sixth list of residuals, discovered additional periodicities of $14\frac{1}{3}$, $19\frac{1}{3}$, and $24\frac{1}{2}$ months. When all had been removed from the data, no other periodicities seemed worth investigation in the residual plot remaining. It is plotted as curve B of figure 1.⁶ The mean of the departures from 1.945 calories in curve B is 0.00189 calorie, or 0.097 percent of the solar constant. Many of the larger departures, which materially raise the mean as just given, occur in

⁶ One disturbing feature will be noted in figure 1, B. Though the year 1947 shows no remarkable eccentricity in curve A, it gives a great slump of $\frac{1}{2}$ percent in curve B. This is strange, for all the periodicities seem to fit the last year's data, including 1946, as well as the earlier years, as we see from figure I, B. One notes, however, that curve A of figure I is almost entirely below I.945 calories in 1947. It may be that the Montezuma values of 1947 are subject to a yet undiscovered error. Further observations of future years

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TABLE 2.-Detailed periodicities

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the months December to February, when the atmospheric conditions at Montezuma are less favorable, and when many days are lost to observation. It cannot be claimed that the periodicities removed are perfectly correct in forms. Hence the final residuals are larger than they should be on this account also. We may conclude that of the variations of solar radiation indicated in figure I,A, and which exceed I percent in range, accidental error of observation contributes less than 2/10 percent, and the periodic variations nearly I percent of the total range.

In table 2 I gave the details of the 14 periodicities in the variation of solar radiation which have been discovered. There may be others of less than 5-months period, some of minor amplitude, and still others exceeding 273 months in period, which our observations have not yet continued long enough to discover. Indeed the large fluctuations of Great Lakes levels occurring at intervals of about 45 and 9I years seem to indicate that the double and quadruple of the master period of 273 months are of very great importance in meteorology. There is also the noted Bruckner period, of about three sunspot cycles, which may also be found eventually in solar-constant values if they continue to be observed for some years longer.